ASTRONOMY 8300 – Fall 2024 Homework Set 1 Answers

1.a. Maxwellian distr: $P(\mathbf{v}_r) = \frac{1}{\sqrt{\pi b}} e^{-(\mathbf{v}_r/b)^2}$ Gaussian profile: $P(\mathbf{v}_r) = \frac{1}{\sqrt{2\pi\sigma}} e^{-\frac{1}{2}(\mathbf{v}_r/\sigma)^2} \rightarrow b = \sqrt{2\sigma}$ $P(0) = \frac{1}{\sqrt{2\pi\sigma}}$ (amplitude of Gaussian) $P(HWHM) = \frac{1}{2}P(0) = \frac{1}{2}\frac{1}{\sqrt{2\pi\sigma}} = \frac{1}{\sqrt{2\pi\sigma}} e^{-\frac{1}{2}(HWHM/\sigma)^2}$ $\frac{1}{2} = e^{-\frac{1}{2}(\mathbf{v}_r/\sigma)^2} \rightarrow \ln(\frac{1}{2}) = -\frac{1}{2}(HWHM/\sigma)^2$ $HWHM = \sqrt{2\ln(2)\sigma} = 1.1774\sigma$ $FWHM = 2HWHM = 2.355\sigma$

1.b.
$$b = \sqrt{\frac{2kT}{m}} = 0.17 \text{ km s}^{-1}$$
 where $m = (28.09)(1.67 \times 10^{-24} \text{ g})$
 $\sigma = \frac{b}{\sqrt{2}} = 0.12 \text{ km s}^{-1}$
FWHM = 2.355 $\sigma = 0.28 \text{ km s}^{-1}$
1.c. In radial velocity: *FWHM*(*LSF*) = $\frac{\Delta\lambda}{\lambda}$ c = 11.9 km s⁻¹ \rightarrow no

1.d. Linear part of the curve of growth for one line (lower limit to column density), full COG for more than one line: FWHM(line) << FWHM(LSF) (lines are unresolved)

1.e. Optical depth method: FWHM(line) > FHWM(LSF) (line is resolved)



2. Absorption-line (F_{λ}/F_c) and optical depth ($\tau = - \ln [F_{\lambda}/F_c]$) profiles:

2.d.
$$N = 1.1298 \times 10^{20} \frac{1}{f\lambda^2} \int \tau_{\lambda} d\lambda = \frac{1}{(0.959)(1260.4)^2} (1.85) = 1.37 \times 10^{14} \text{ cm}^{-2}$$

3. Use oscillator strengths (f) from Morton et al. (1991). Determine W_{λ}/λ and Nf λ as a function of τ_0 for different values of b to get curves of growth. Note for the equation:

$$\tau_{0} = \frac{N_{j}s\lambda_{jk}}{\sqrt{\pi}b} = \frac{1.497 \ x \ 10^{-2}}{b} N_{j}\lambda_{jk}f_{jk},$$

the parameters are in cgs units, so b should be in cm/sec (see Spitzer, Ch. 3). Assume different columns to get Nf λ for data points and overplot on curves of growth: Answer: N(Si II) = 4 (± 1) x 10¹³ cm⁻², b = 12 (± 2) km s⁻¹



Propagation of Errors (Bevington 1969):

$$y = \log\left(\frac{W_{\lambda}}{\lambda}\right) = \frac{1}{2.3} \ln\left(\frac{W_{\lambda}}{\lambda}\right)$$
$$\sigma_{y} = \frac{1}{2.3} \left(\frac{\sigma_{W_{\lambda}}}{W_{\lambda}}\right)$$