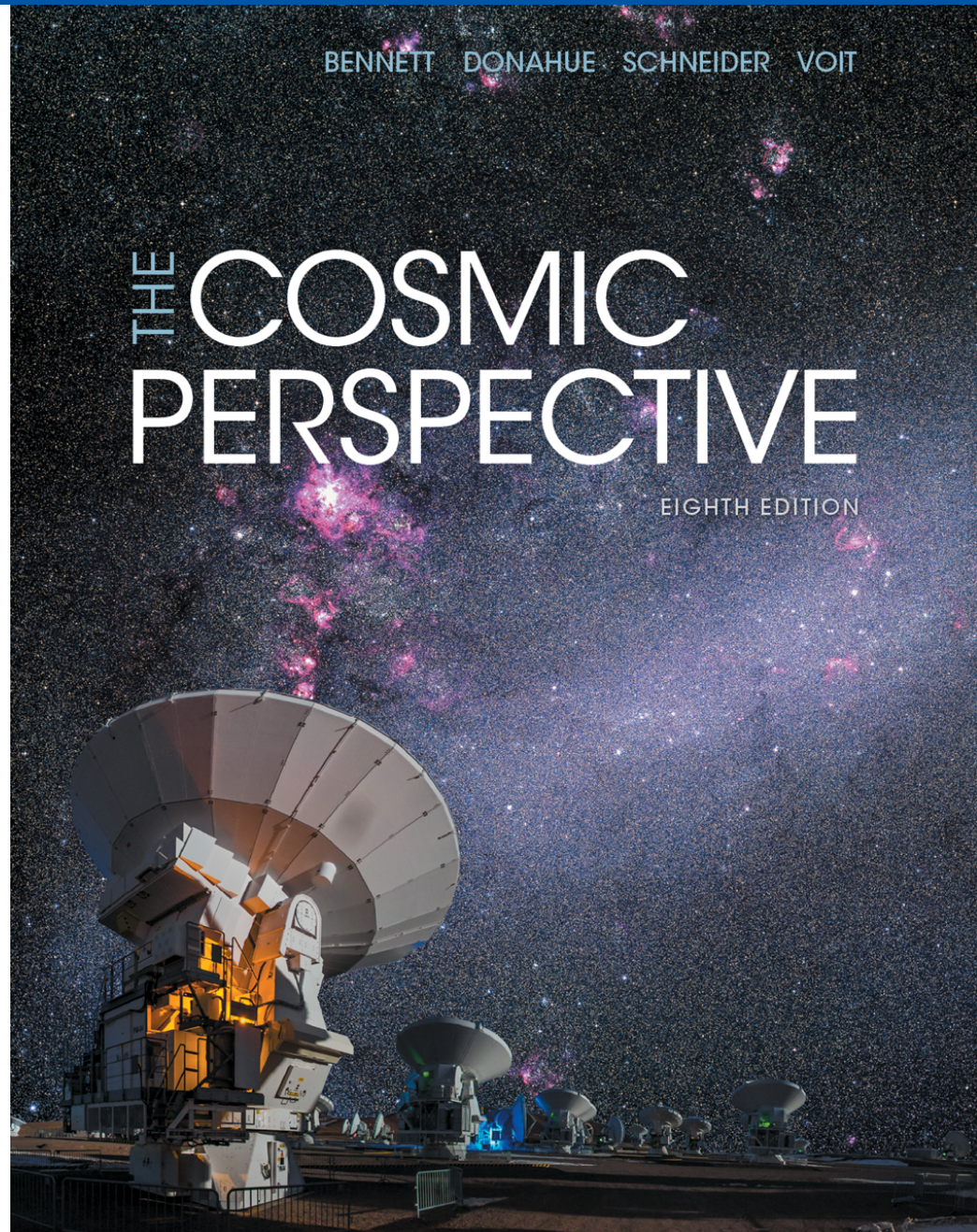


Chapter 4: Making Sense of the Universe: Understanding Motion, Energy, and Gravity



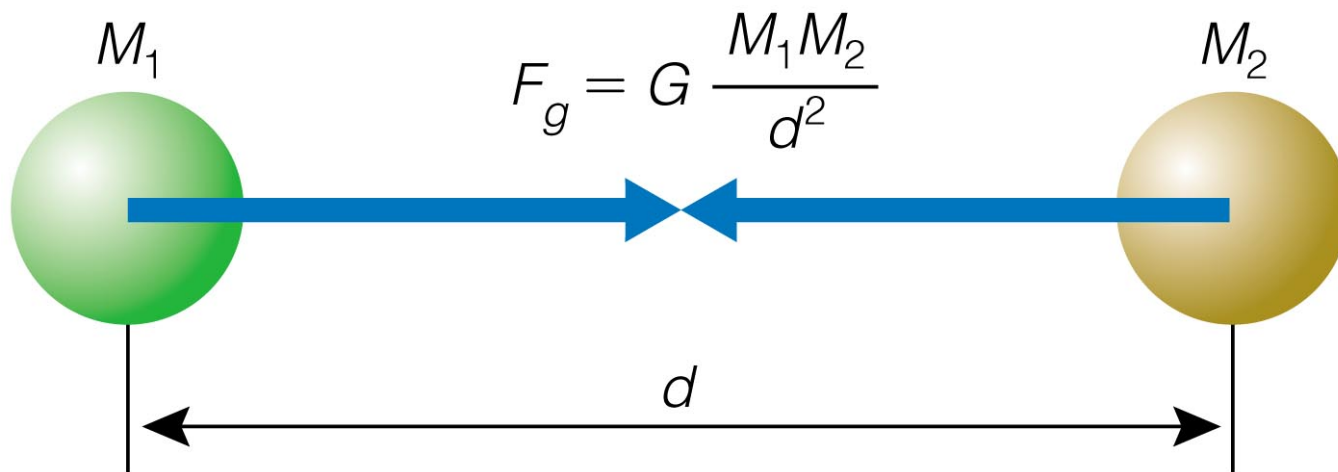
4.4 The Universal Law of Gravitation

- Our goals for learning:
 - **What determines the strength of gravity?**
 - **How does Newton's law of gravity extend Kepler's laws?**

What determines the strength of gravity?

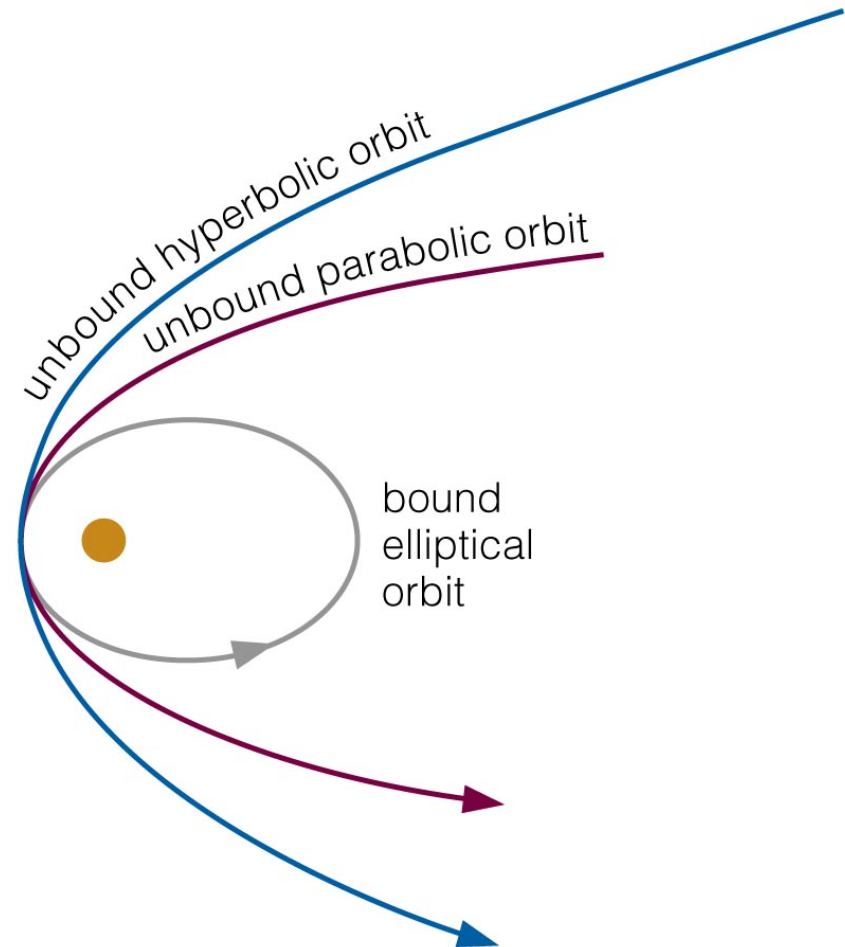
The universal law of gravitation:

1. Every mass attracts every other mass – gravitational forces are only attractive
2. Attraction is *directly* proportional to the product of the interacting masses.
3. Attraction is *inversely* proportional to the *square* of the distance between their centers.



How does Newton's law of gravity extend Kepler's laws?

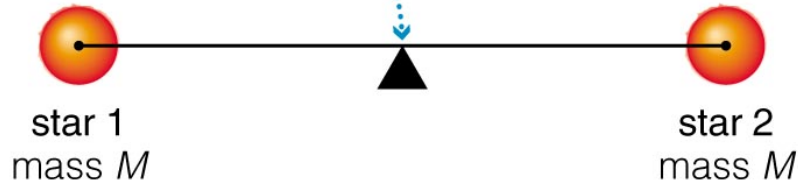
- Kepler's laws apply to all orbiting objects, not just planets.
- Ellipses are not the only orbital paths. Orbits can be:
 - bounded (ellipses)
 - unbounded
 - parabola
 - hyperbola



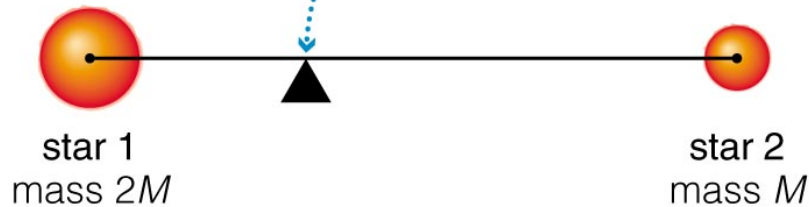
a Orbits allowed by the law of gravity.

Center of Mass

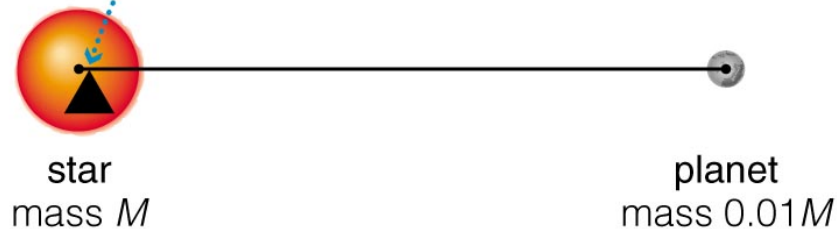
For two stars of equal mass: The center of mass lies halfway between them.



For two stars with different masses: The center of mass lies closer to the more massive one.



The star is so much more massive than the planet that the center of mass lies inside the star.



- Because of momentum conservation, orbiting objects orbit around their center of mass.

Newton and Kepler's Third Law

- Newton's laws of gravity and motion showed that the relationship between the *orbital period* and *average orbital distance* of a system tells us the *total mass* of the system.
- Examples:
 - Earth's orbital period (1 year) and average distance (1 AU) tell us the Sun's mass.
 - Orbital period and distance of a satellite from Earth tell us Earth's mass.
 - Orbital period and distance of a moon of Jupiter tell us Jupiter's mass.

Newton's Version of Kepler's Third Law

$$p^2 = \frac{4\pi^2}{G(M_1 + M_2)} a^3 \quad \text{OR} \quad M_1 + M_2 = \frac{4\pi^2}{G} \frac{a^3}{p^2}$$

- p = orbital period
- a = average orbital distance (between centers)
- $(M_1 + M_2)$ = sum of object masses

What have we learned?

- **What determines the strength of gravity force?**
 - Directly proportional to the *product* of the masses ($M \times m$)
 - *Inversely* proportional to the *square* of the separation
- **How does Newton's law of gravity allow us to extend Kepler's laws?**
 - Applies to other objects, not just planets
 - Includes unbound orbit shapes: parabola, hyperbola
 - Can be used to measure mass of orbiting systems

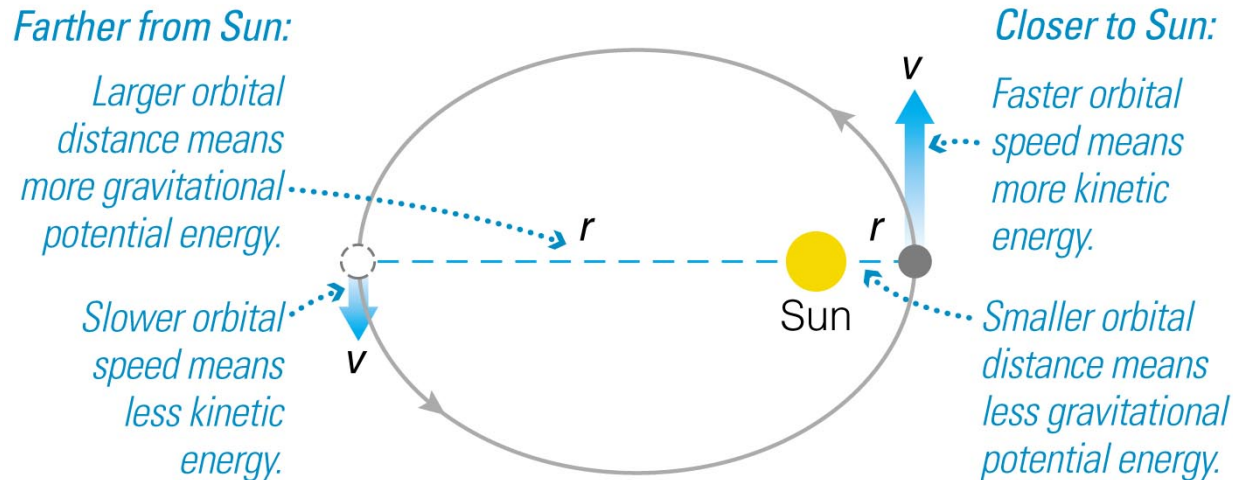
4.5 Orbits, Tides, and the Acceleration of Gravity

- Our goals for learning:
 - **How do gravity and energy together allow us to understand orbits?**
 - **How does gravity cause tides?**
 - **Why do all objects fall at the same rate?**

How do gravity and energy together allow us to understand orbits?

- Total orbital energy (gravitational + kinetic) stays constant if there is no external force.
- Orbits cannot change spontaneously.

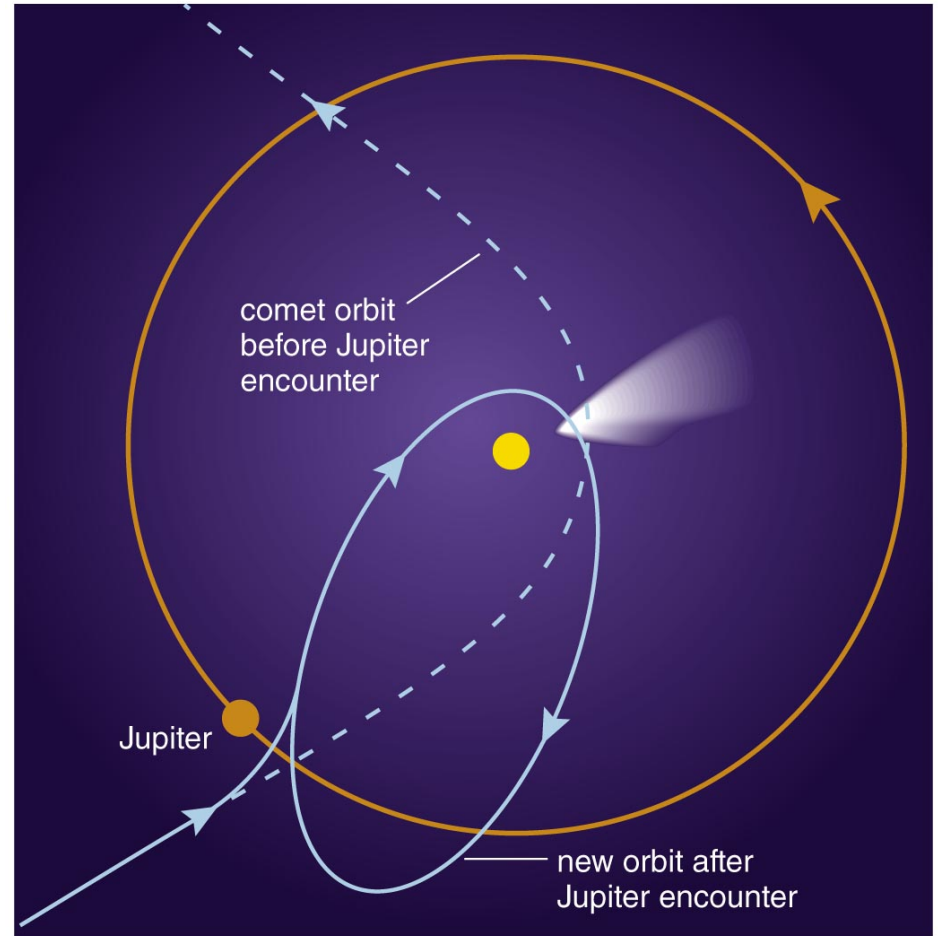
Total orbital energy = gravitational potential energy + kinetic energy



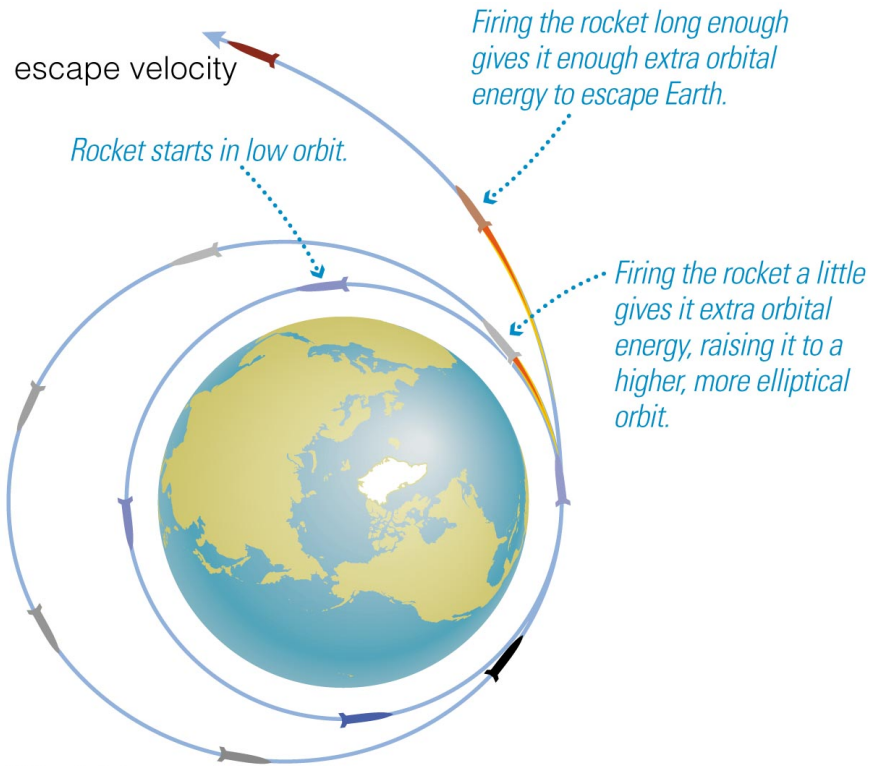
Total orbital energy stays constant.

Changing an Orbit

- So what can make an object gain or lose orbital energy?
 - Friction or atmospheric drag
 - A gravitational encounter



Escape Velocity

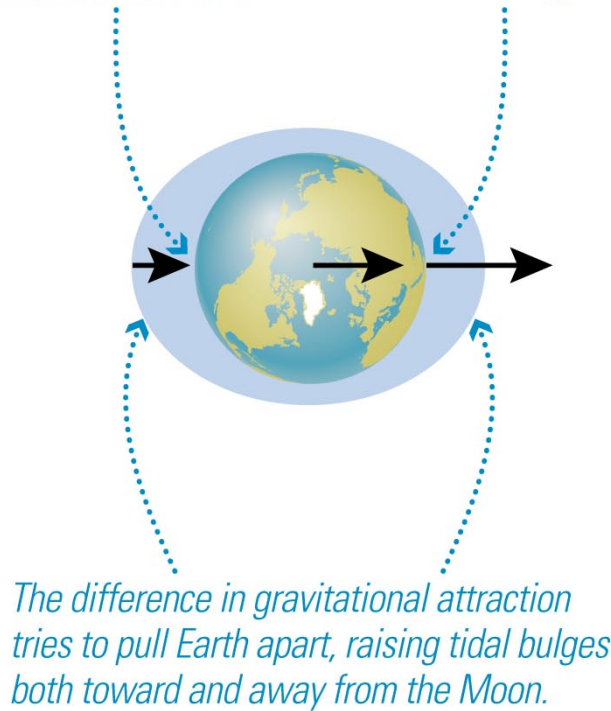


- If an object gains enough orbital energy, it may escape (change from a bound to unbound orbit).
- **Escape velocity** from Earth ≈ 11 km/s from sea level (about 40,000 km/hr [$\sim 25,000$ miles per hour])

How does gravity cause tides?

The gravitational attraction to the Moon is weakest here . . .

. . . and strongest here.



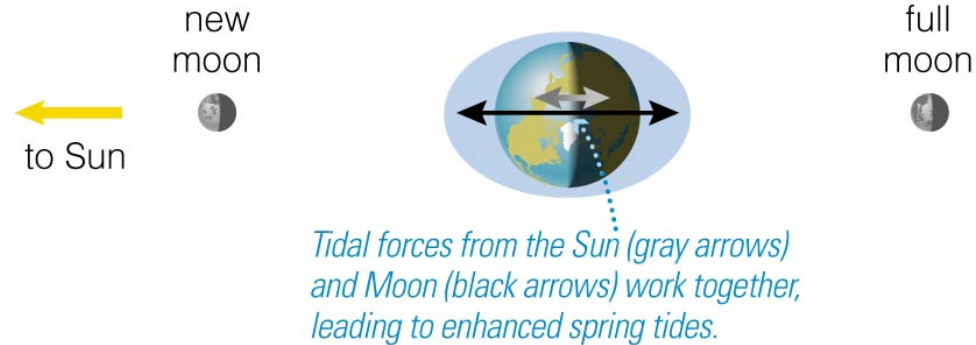
Not to scale!

- Moon's gravity pulls harder on near side of Earth than on far side.
- Difference in Moon's gravitational pull stretches Earth.

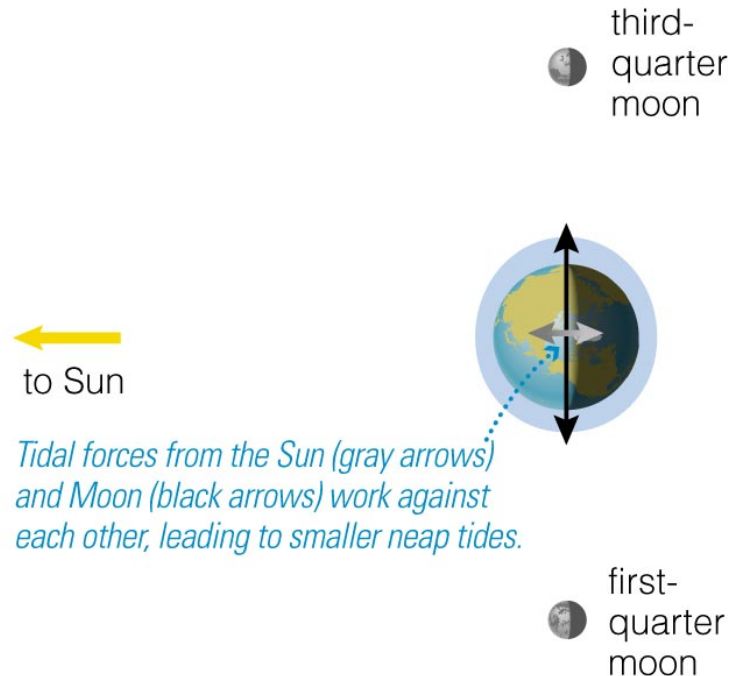
Tides and Phases

- The Sun also has a small tidal effect on Earth.
- Size of tides thus depends on phase of Moon.

Spring tides occur at new moon and full moon:



Neap tides occur at first- and third-quarter moon:



Tidal Friction

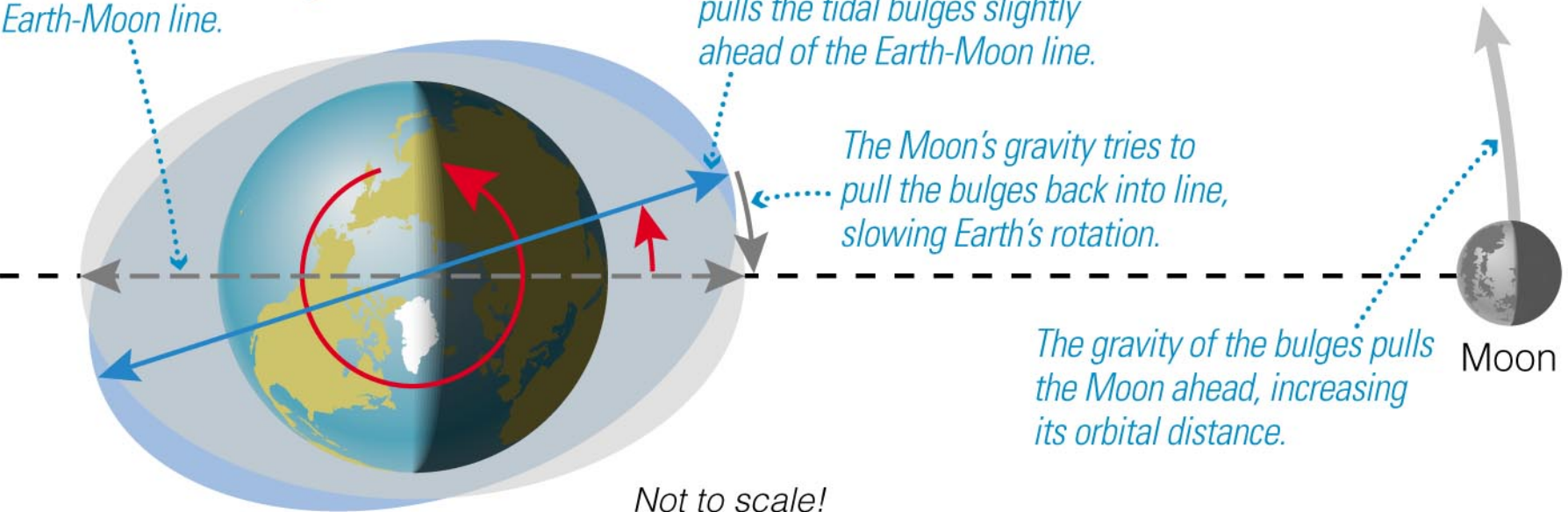
If Earth didn't rotate, tidal bulges would be oriented along the Earth-Moon line.

Friction with the rotating Earth pulls the tidal bulges slightly ahead of the Earth-Moon line.

The Moon's gravity tries to pull the bulges back into line, slowing Earth's rotation.

The gravity of the bulges pulls the Moon ahead, increasing its orbital distance.

Not to scale!



- Tidal friction gradually slows Earth's rotation (and makes the Moon get farther from Earth).
- The Moon once orbited faster (or slower); tidal friction caused it to "lock" in synchronous rotation.

Why do all objects fall at the same rate?

$$a_{\text{rock}} = \frac{F_g}{M_{\text{rock}}} \qquad F_g = G \frac{M_{\text{Earth}} M_{\text{rock}}}{R_{\text{Earth}}^2}$$
$$a_{\text{rock}} = G \frac{M_{\text{Earth}} \cancel{M_{\text{rock}}}}{R_{\text{Earth}}^2 \cancel{M_{\text{rock}}}} = G \frac{M_{\text{Earth}}}{R_{\text{Earth}}^2}$$

- The gravitational acceleration of an object like a rock does not depend on its mass because M_{rock} in the equation for acceleration cancels M_{rock} in the equation for gravitational force.
- This initially surprising result was fully explained by Einstein, in his general relativity theory.

Proof-of-concept experiment

Feather & Ball
Sealing Drop &

What have we learned?

- **How do gravity and energy together allow us to understand orbits?**
 - Change in total energy is needed to change orbit
 - Add enough energy (escape velocity) and object leaves.
- **How does gravity cause tides?**
 - The Moon's gravity stretches Earth and its oceans.
- **Why do all objects fall at the same rate?**
 - Mass of object in Newton's second law exactly cancels mass in law of gravitation.