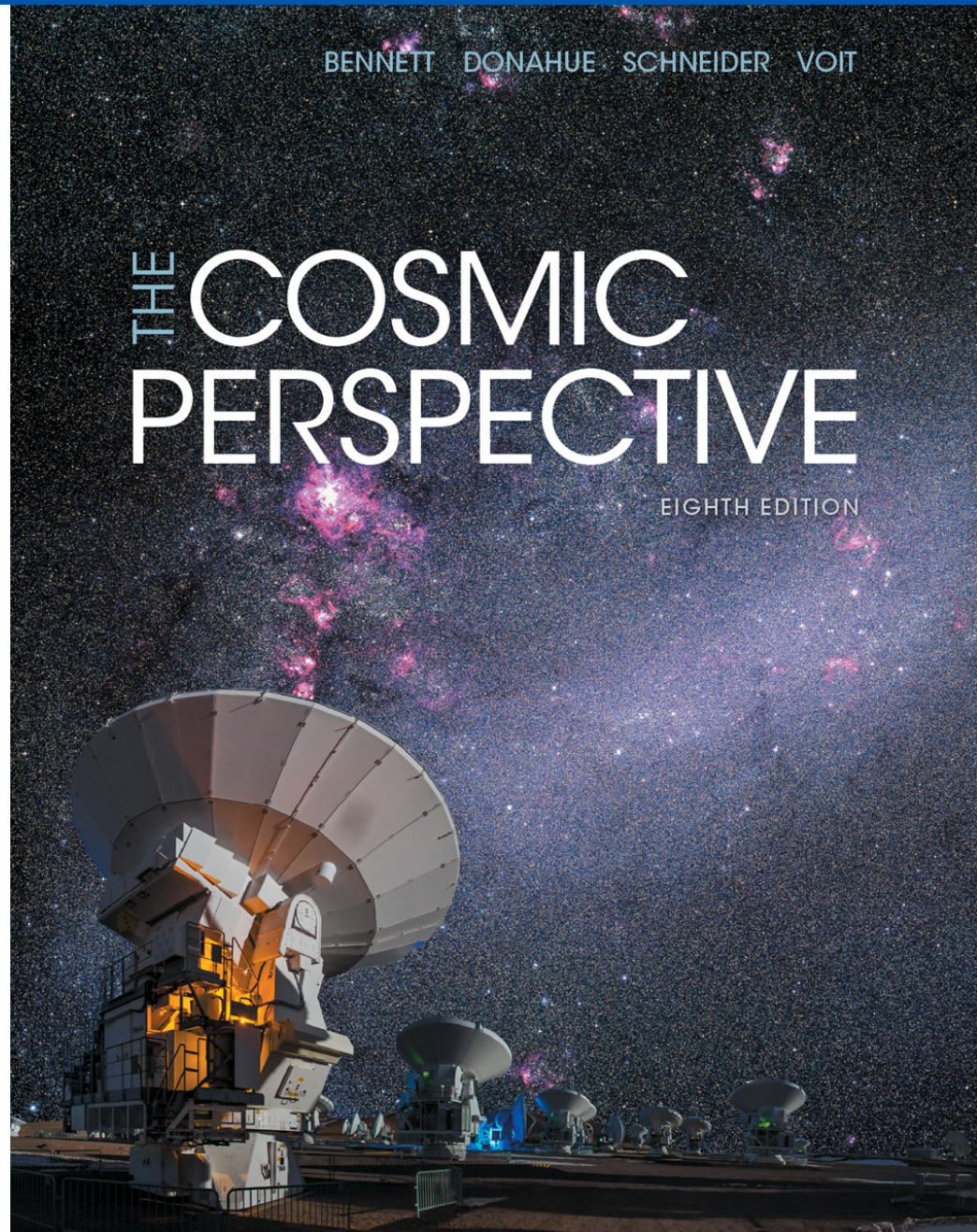


## Chapter S1: Celestial Timekeeping and Navigation



# S1.1 Astronomical Time Periods

- How do we define the day, month, year, and planetary periods?
- How do we tell the time of day?
- When and why do we have leap years?

# How much does Earth rotate around its axis during one solar day?

- a) 360 degrees
- b) about 359 degrees
- c) about 361 degrees
- d) about 365 degrees
- e) Earth does not rotate.

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- a) sidereal month.
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- c) solar month.
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# The difference between a tropical year and a sidereal year is due to the

- a) combination of Earth's rotation and Earth's orbital motion.
- b) mismatch of the exact length of a year and 365 days.
- c) tug of the Moon on Earth as it orbits the Sun.
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# When is an object at opposition?

- a) when the object is on the opposite side of the Earth as the Sun
- b) when the object is on the opposite side of the Sun as the Earth
- c) when the Earth and the Sun are on opposite sides of the object
- d) when the object's motion relative to the stars changes direction and becomes retrograde

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- d) when the object's motion relative to the stars changes direction and becomes retrograde

Which of the following statements about mean solar time and apparent solar time is not true?

- a) Mean solar time is based on the average length of the solar day.
- b) Apparent solar time measures noon when the Sun is on the meridian.
- c) Mean solar time is sometimes ahead of apparent solar time and sometimes behind apparent solar time.
- d) Mean solar time and apparent solar time are always within about 4 minutes of each other.

Which of the following statements about mean solar time and apparent solar time is *not* true?

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# Leap years account for the

- a) combination of Earth's rotation and Earth's orbital motion.
- b) mismatch of the exact length of a year and 365 days.
- c) difference between sidereal years and tropical years.
- d) changing length of the year from one year to the next.

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# S1.2 Celestial Coordinates and Motion in the Sky

- How do we locate objects on the celestial sphere?
- How do stars move through the local sky?
- How does the Sun move through the local sky?

# How does the position of the Sun on the celestial sphere change over the course of a year?

- a) Its right ascension changes and its declination remains constant.
- b) Its right ascension remains constant and its declination changes.
- c) Its celestial coordinates remain fixed.
- d) Both its right ascension and declination change.



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# The declination of an object tells us its angular distance

- a) above the horizon.
- b) from the meridian.
- c) from the prime meridian.
- d) from the celestial equator.
- e) from the ecliptic.

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- d) from the celestial equator.**
- e) from the ecliptic.

At what altitude does the celestial equator cross your meridian?

- a) 90 degrees
- b) 90 degrees minus your latitude
- c) your latitude
- d) 0 degrees
- e) The celestial equator does not cross the meridian.

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# What is the arctic circle?

- a) It is the region of the arctic north of all land masses.
- b) It is the path that the Sun makes in the sky over the course of a year when viewed from the north pole.
- c) It is the latitude in the northern hemisphere north of which the Sun doesn't set for at least one full day.
- d) It is the latitude in the northern hemisphere north of which the Sun doesn't rise for at least one full day.
- e) C and D

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# When is the Sun highest in the sky at noon for someone who lives on the equator?

- a) on the equinoxes (March 21 and September 21)
- b) on the solstices (June 21 and December 21)
- c) on the winter solstice (December 21)
- d) It is the same altitude at noon on all days.
- e) when the Moon is in the 7<sup>th</sup> house and Jupiter aligns with Mars



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During which period of time is the Sun continuously below the horizon for someone at the south pole?

- a) from March 21 to September 21
- b) from June 21 to December 21
- c) from September 21 to March 21
- d) from December 21 to June 21

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Stars cross the local meridian at an altitude equal to their

- a) declination.
- b) declination plus your latitude.
- c) declination minus your latitude.
- d) none of the above

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# S1.3 Principles of Celestial Navigation

- How can you determine your latitude?
- How can you determine your longitude?

# The altitude of the North Celestial Pole in your local sky

- a) is equal to your latitude.
- b) is equal to 90 degrees minus your latitude.
- c) is equal to 90 degrees.
- d) depends on the time of day.
- e) depends on the season.

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# To measure latitude, one needs a way to measure

- a) angles in the sky.
- b) time very accurately.
- c) A and B.

# To measure latitude, one needs a way to measure

**a) angles in the sky.**

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c) A and B.

# To measure longitude, one needs

- a) to measure angles in the sky.
- b) to have a very accurate clock.
- c) to have a way to communicate instantly with someone at a different, known longitude.
- d) all of the above.
- e) B or C.

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- a) to measure angles in the sky.
- b) to have a very accurate clock.
- c) to have a way to communicate instantly with someone at a different, known longitude.
- d) all of the above.
- e) B or C.**

If a star crosses the meridian as seen from Greenwich, England, at midnight, when does it cross the meridian for someone at a longitude of 90 degrees West?

- a) midnight
- b) 6 a.m.
- c) 6 p.m.
- d) 9 a.m.
- e) 9 p.m.

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