



Earth, the “Goldilocks” Planet

What We Will Learn Today

- What are planet Earth's features?
- Why does our atmosphere have different layers?
- Why is our sky blue?
- What creates wind and weather?
- How will our climate change over the long run?
- How is Earth's Atmosphere unique in the solar system?

Planet Earth Details

Property	Value
Semi-major axis	149.6 million km (1.00 AU)
Size (radius)	6,378 km (1.00 R _{Earth})
Mass	5.97 x 10 ²⁴ kg (1.00 M _{Earth})
Average density	5.52 g/cc
Composition	Rocks, metals
Average surface temperature	290 K
Moons	1
Orbital period	365.242 days (tropical)
Rotation period (sidereal)	23h, 56m (sidereal)
Axis tilt	23.5°
Orbital inclination	0.00°
Orbital eccentricity	0.007

Interaction of Sunlight With Atmosphere

- X rays from the Sun ionize atoms high in the atmosphere
 - absorbed in the process
- UV photons disassociate loosely bound molecules such as ozone (O_3)
- Visible light is scattered
 - Gives us a blue sky
- IR light from the planet is trapped
 - Greenhouse effect

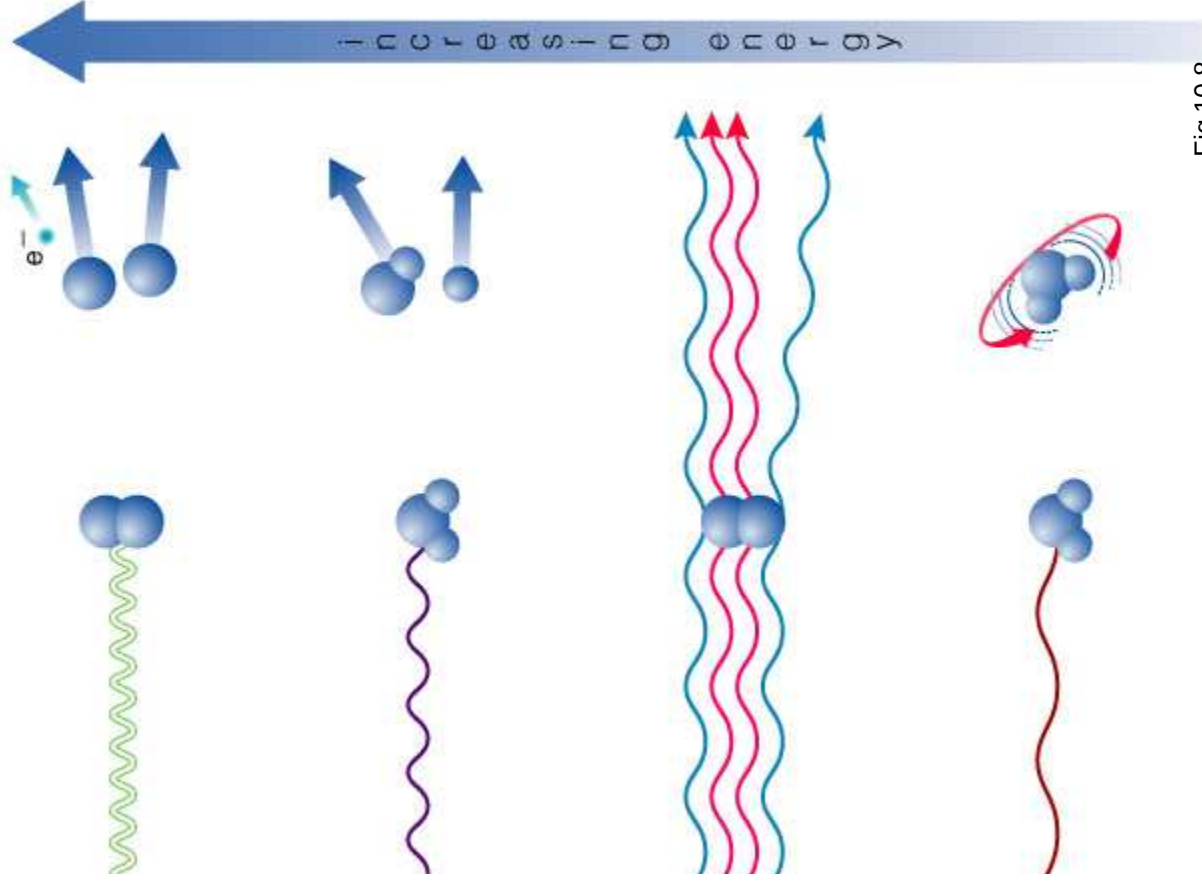


Fig 10.8

Structure of Our Atmosphere

- Troposphere
 - Heated by greenhouse gases
 - Convection is important
 - Tropos = turning
 - Temperature drops with altitude (3.5 °F per 1000 ft)
- Stratosphere
 - Heated by UV light
 - Temperature rises with altitude, then drops
 - No convection
 - Stratified, stable layers
 - No weather
 - Unique to Earth among terrestrial worlds

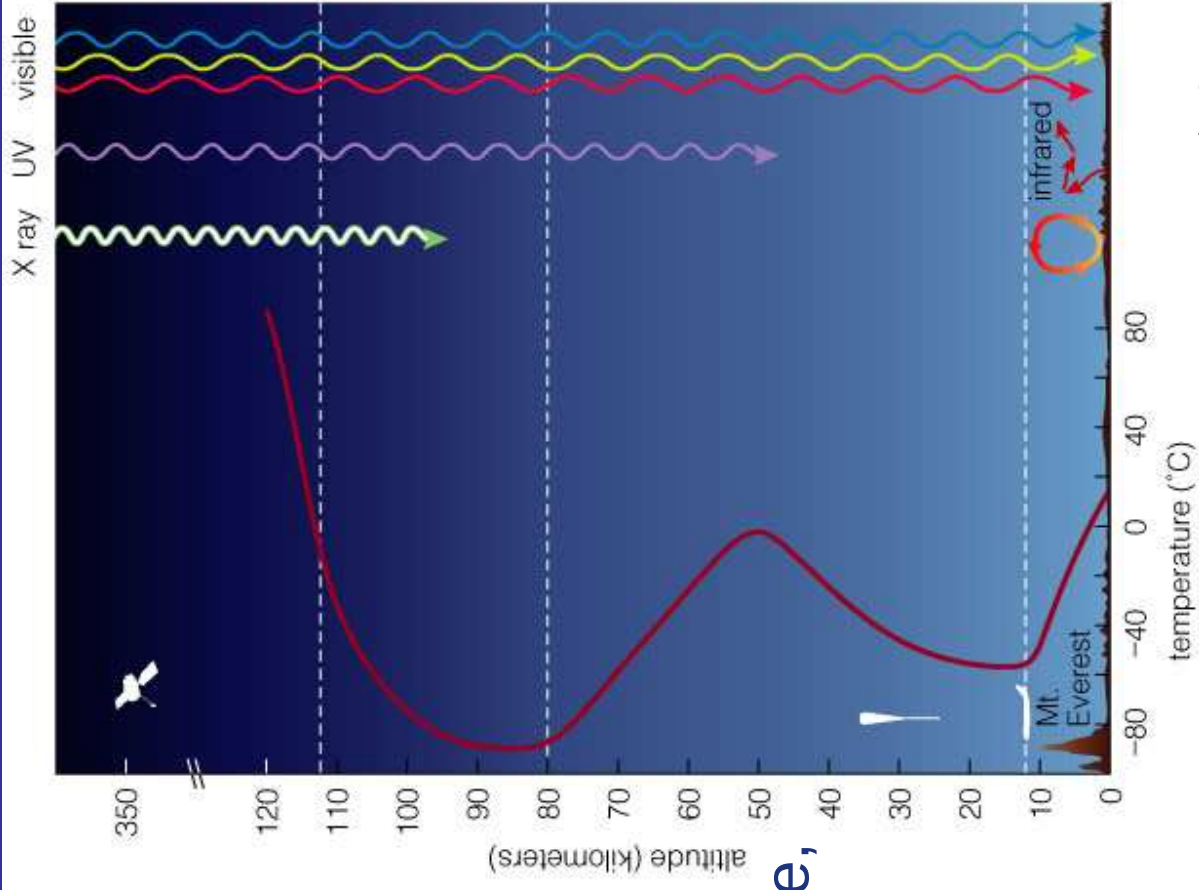


Fig 10.7

Structure of Our Atmosphere

- Thermosphere
 - Heated by X-rays from the Sun
 - Temperature rises with altitude
 - Would not feel hot!
 - Low density
 - Partly ionized (ionosphere)
 - Reflects radio waves back to Earth
- Exosphere
 - Atmosphere fades out
 - Black sky, weightless

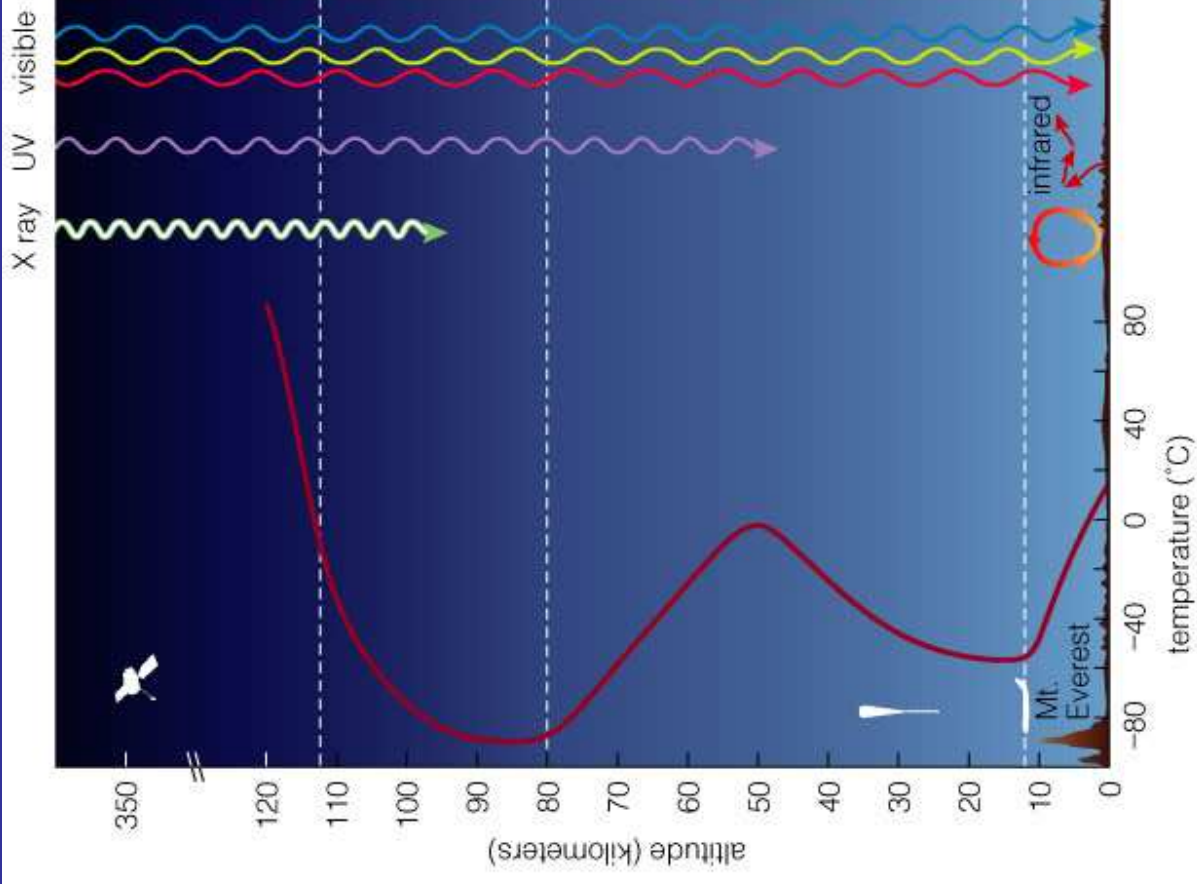


Fig 10.7

Why Is the Sky Blue?

- Sun emits most of its light in visible (ROYGBIV)
- Shorter wavelengths interact more with matter
- Blue light is scattered off air molecules
- Most of the red light reaches the surface
- Results
 - Blue sky
 - Red sun during rise/set
 - Bright sky (no stars during the day)

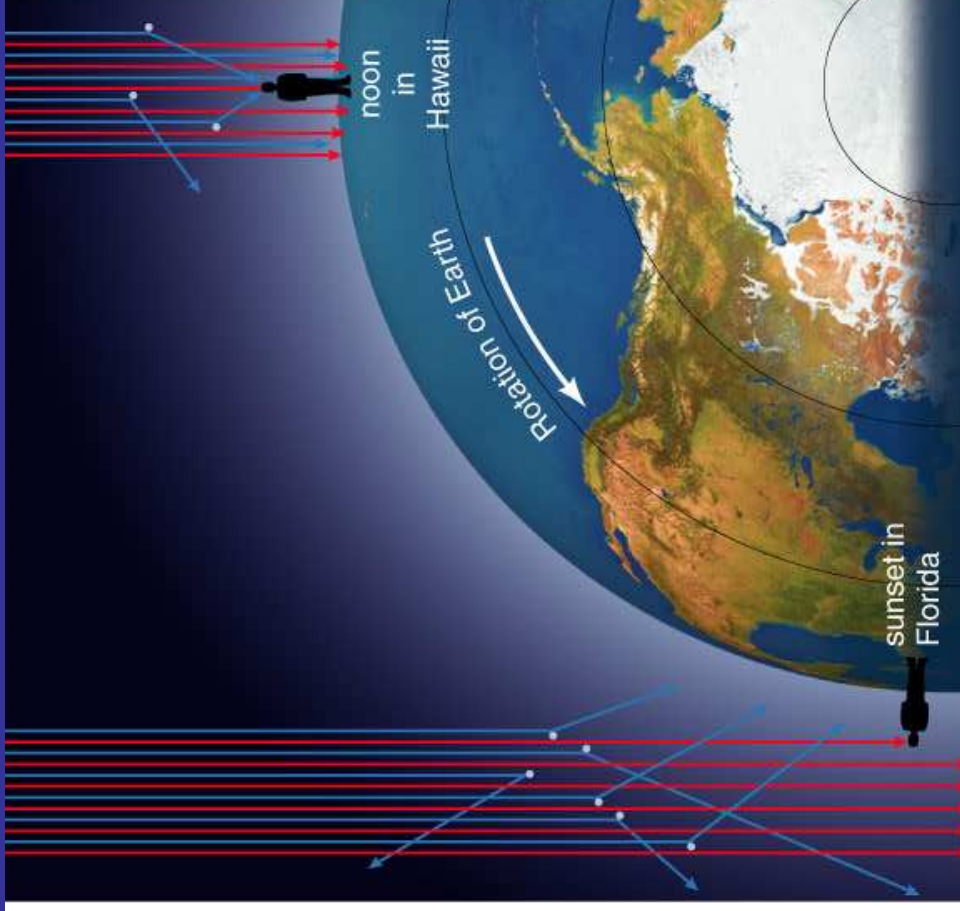


Fig 10.9

Protective Magnetosphere

- Solar wind: High velocity charged particles from the Sun
- Earth's magnetic field protects the planet
- Creates beautiful aurorae
 - Interaction of charged particles with atmosphere
- Van Allen belts
 - Charged particles inside magnetosphere moving along magnetic field lines

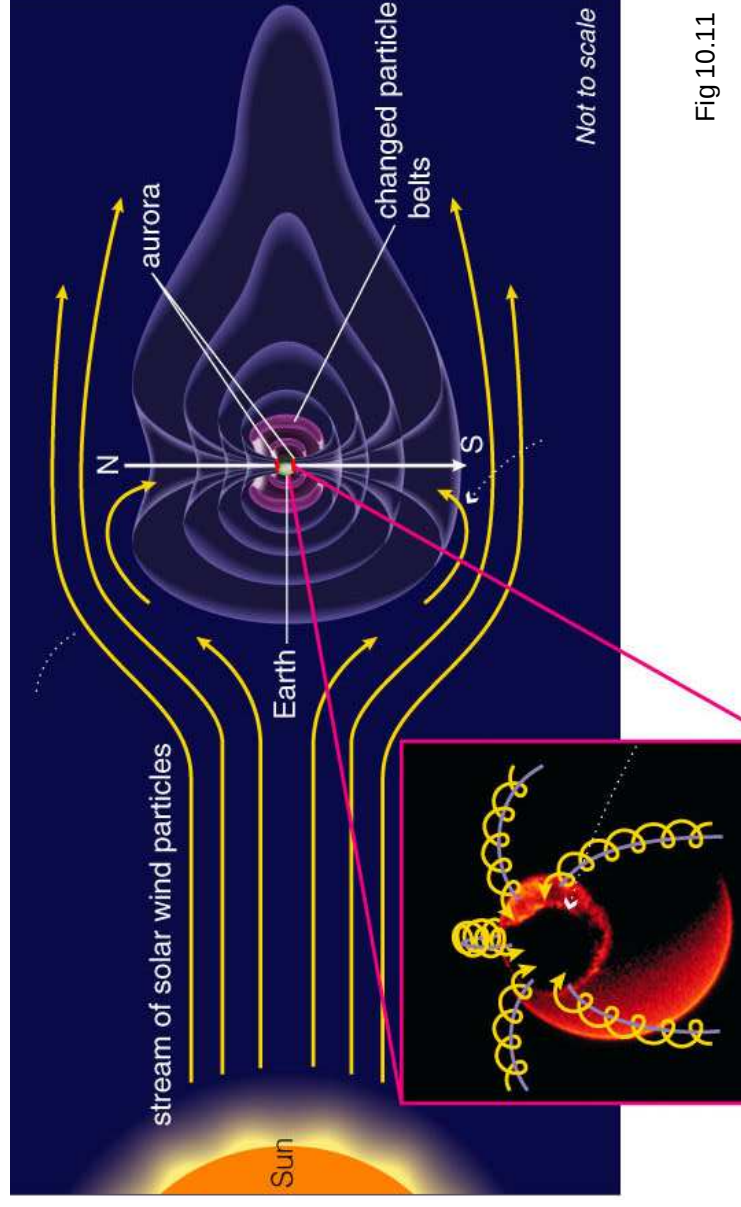


Fig 10.11

Weather Patterns: Wind

- Heating causes circulation cells
 - Warm air rises at the equator
 - Flows towards the pole, where it descends
 - Carries heat from equator to pole
 - Remember, Venus is very efficient at this
- Coriolis Effect
 - Earth's rotation causes circular motion of wind
 - Breaks down the circulation cell into three cells per hemisphere
 - Unique to Earth among terrestrial worlds
- Resulting surface winds
 - East to West near equator and poles
 - West to East in mid latitudes

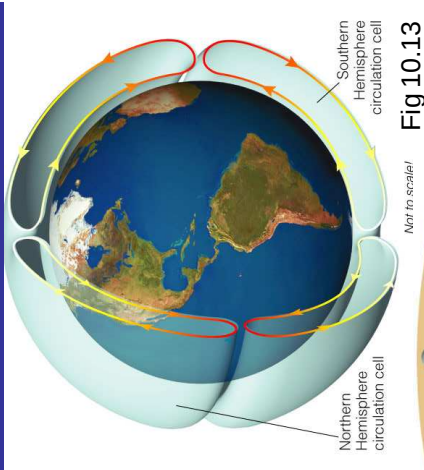


Fig 10.13

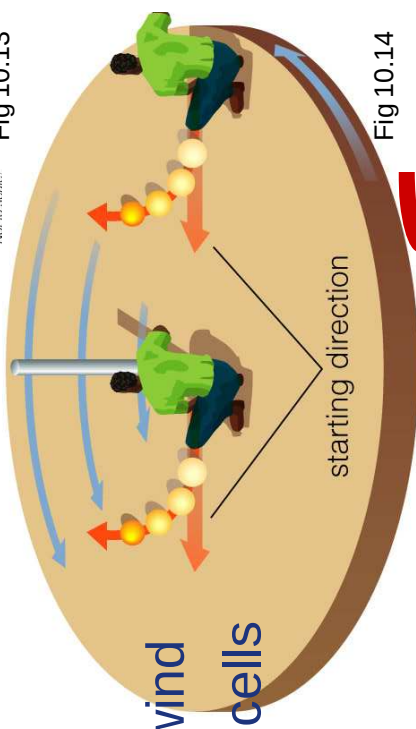


Fig 10.14



Fig 10.15

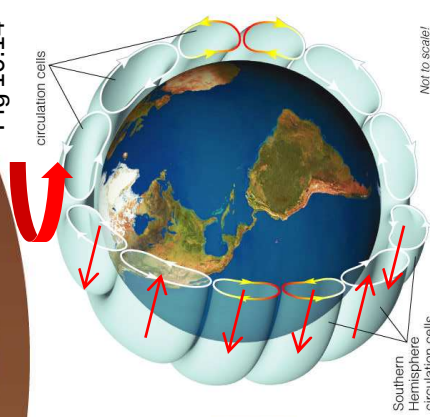


Fig 10.16

Long-term Climate Changes

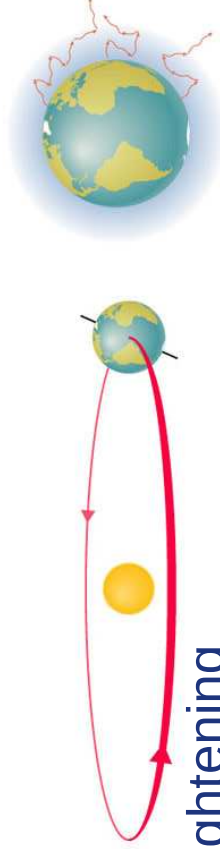
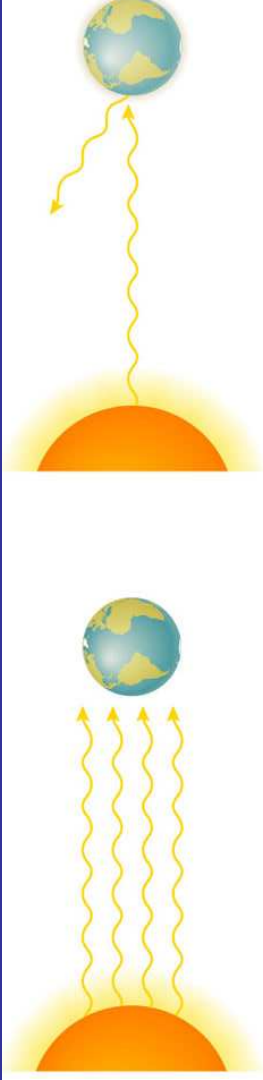


Fig 10.18

- Solar brightening
 - Sun is 30% more luminous now compared to when it was born
 - In a billion years, will be hot enough to evaporate our oceans (runaway greenhouse effect)
 - Earth will become a molten rock ball in a few billion years
- Changes in axial tilt
 - Varies over 22° – 25° over tens of thousands of years
 - Bigger factor for Mars
- Changes in reflectivity
 - Clouds and ice are better reflectors
- Changes in greenhouse gas abundance

Ice Ages on Earth

- CO₂ cycle is 400,000 years long
- Gives room for climate changes such as ice ages
 - 100,000 year cycles
 - ~ 85,000 years of ice age
 - ~ 15,000 years of interglacial periods
- Snowball Earth
 - 750 – 580 million years ago, the entire Earth was frozen over
 - Rescued by plate tectonics and volcanism

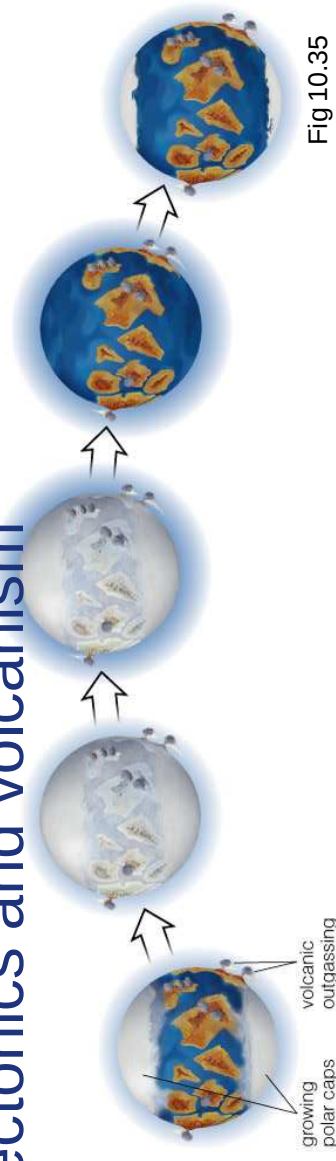
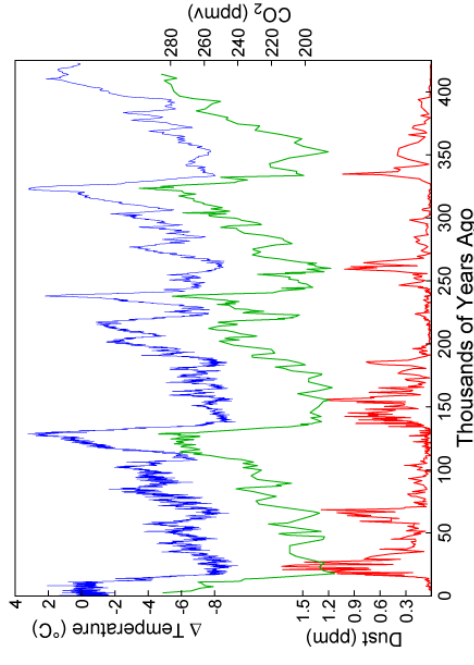


Fig 10.35

Earth's Regulated Greenhouse

- CO₂ and H₂O are efficient greenhouse gases
- CO₂ has a self-regulating effect!
 - If volcanoes put excess CO₂ in atmosphere
 - Temperatures rise
 - Convection & rainfall will increase
 - Rain traps more CO₂ from atmosphere into ocean and rocks
 - Earth cools back down
 - Dominant regulator on Earth
- H₂O has a runaway effect
 - If increased heat evaporates more of the oceans
 - Water vapor in atmosphere traps more heat
 - This increases temperature and evaporation
 - Increases greenhouse gas content in atmosphere
 - Runaway greenhouse effect
 - Dominant process in the young Venus

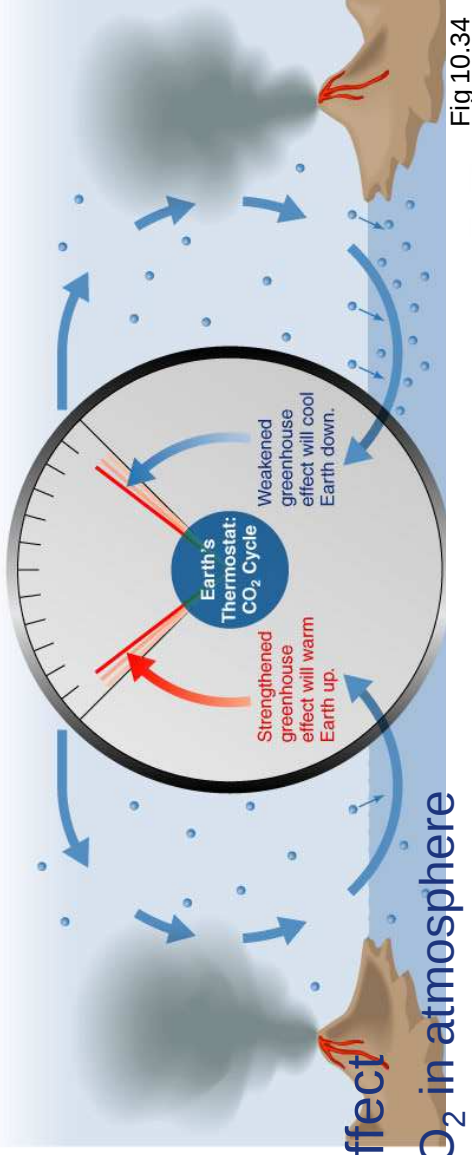


Fig 10.34



Earth's Perfect Environment

- Retained most of the outgassed water
 - Moderate temperatures allowed condensation
 - Create oceans of water on the surface
- Locked away most of the CO₂
 - Dissolved in oceans (60x atmosphere) and locked in carbonate rocks (170,000x atmosphere)
- Nitrogen becomes primary constituent (77%)
 - CO₂ and H₂O locked removed from the air
- Oxygen (21%) level is maintained by life
 - Very reactive, will be lost if not replenished
 - Creates the UV absorbent ozone layer

Human Influence on our Climate

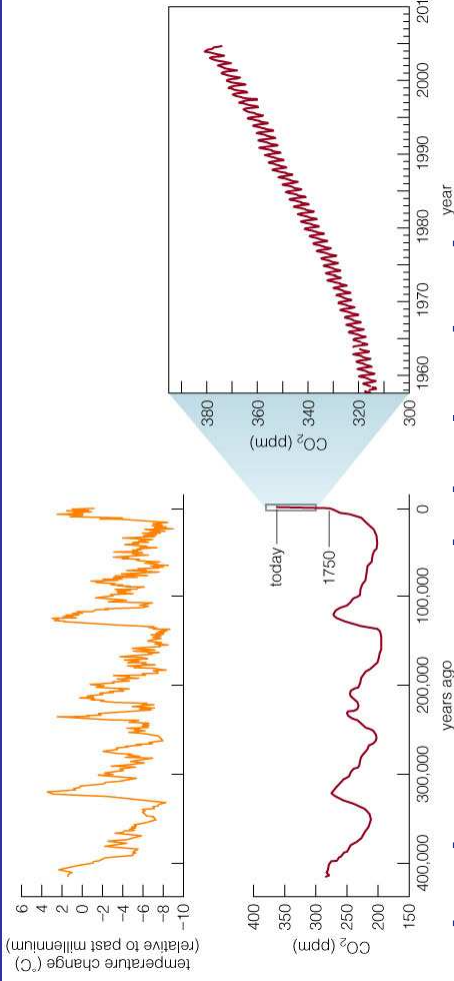


Fig 10.36

- CFC emission created “hole” in the ozone layer
- Extinction rate of other species accelerate due to our dominance
- Global warming
 - A complex process
 - We don’t fully understand it
 - But we are undoubtedly altering the natural process
- Potential consequence of global warming
 - Largely unknown, but could include:
 - Increased storms, rising ocean levels, what else?

Unique Features of Earth's Atmosphere

- Not too hot, not too cold!
 - Liquid water on surface, water vapor in atmosphere
- Oxygen in the atmosphere
 - Maintained by life on Earth
- Stratosphere
 - Efficient UV absorbers in atmosphere
- Strong magnetosphere
 - Protective blanket from solar wind particles
- Weather driven by Coriolis effect
 - Large size + fast rotation
- Self-regulating greenhouse effect
- Artificial effects on atmosphere